

Estatística quântica

Já vimos anteriormente desvios da estatística clássica através do englobamento de níveis que deixam de participar do teorema de equipartição de energia:

A partir de agora vamos estudar o papel da indistinguibilidade das partículas quânticas no desvio da estatística clássica

Sabemos que o estado físico de N partículas é descrito por $\psi(x_1, x_2, \dots, x_N)$ onde $x_i = \vec{x}_i, b_i$

e que $|\psi(x_1, \dots, x_i, \dots, x_j, \dots, x_N)|^2 = |\psi(x_1, \dots, x_j, \dots, x_i, \dots, x_N)|^2$

$$\Rightarrow \psi(x_1, \dots, x_i, \dots, x_j, \dots, x_N) = e^{i\phi_{ij}} \psi(x_1, \dots, x_j, \dots, x_i, \dots, x_N)$$

Trocando novamente $i \leftrightarrow j$

$$\psi(x_1, \dots, x_i, \dots, x_j, \dots, x_N) = e^{2i\phi_{ij}} \psi(x_1, \dots, x_i, \dots, x_j, \dots, x_N)$$

$$\Rightarrow e^{2i\phi_{ij}} = 1 \Rightarrow 2i\phi_{ij} = 2n\pi \Rightarrow \phi_{ij} = n\pi$$

$$\text{onde } n=0,1 \Rightarrow e^{i\phi_{ij}} = \pm 1 \quad \left. \begin{array}{l} +1 \rightarrow \text{bósons} \\ -1 \rightarrow \text{férmions} \end{array} \right\}$$

$$\psi(x_1, \dots, x_j, \dots, x_N, t) = \sum_{E_1, \dots, E_N} C(E_1, \dots, E_j, \dots, E_N, t) \psi_{E_1}(x_1) \dots \psi_{E_j}(x_j) \dots \psi_{E_N}(x_N)$$

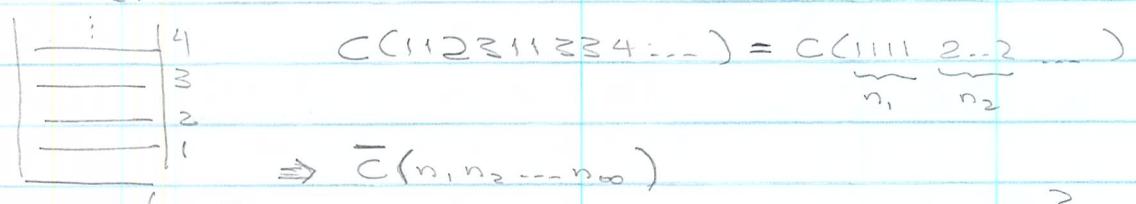
Base: $\{\psi_{E_1}(x_1) \dots \psi_{E_j}(x_j) \dots \psi_{E_N}(x_N)\}$
 \rightarrow auto estados de sistema não interagente

$$\Rightarrow \psi(x_1, \dots, x_j, \dots, x_N, t) = \pm \sum_{E_1, \dots, E_N} C(E_1, \dots, E_j, \dots, E_N, t) \psi_{E_1}(x_1) \psi_{E_j}(x_j) \dots \psi_{E_N}(x_N)$$

$$= \pm \sum_{E_1, \dots, E_N} C(E_1, \dots, E_j, \dots, E_N, t) \psi_{E_1}(x_1) \dots \psi_{E_j}(x_j) \dots \psi_{E_N}(x_N)$$

$$\Rightarrow C(E_1, \dots, E_j, \dots, E_N) = \pm C(E_1, \dots, E_j, \dots, E_N)$$

Bósons:



Como $\int |\psi|^2 dx_1 \dots dx_N = 1 \Rightarrow \sum_{E_1, \dots, E_N} |C(E_1, \dots, E_N)|^2 = 1$

$$\Rightarrow \sum_{n_1, \dots, n_\infty} |\bar{C}(n_1, \dots, n_\infty)|^2 \sum_{\substack{E_1, \dots, E_N \\ (n_1, n_2, \dots, n_\infty)}} 1 = 1$$

$$\Rightarrow \sum_{n_1, \dots, n_\infty} |\bar{C}(n_1, \dots, n_\infty)|^2 \frac{N!}{n_1! n_2! \dots n_\infty!} = 1$$

$\sum n_i = N$

$$f(n_1, \dots, n_\infty) \equiv \sqrt{\frac{N!}{n_1! \dots n_\infty!}} \bar{C}(n_1, \dots, n_\infty)$$

$$\sum_{n_1, \dots, n_\infty} |f|^2 = 1$$

$$\Rightarrow \psi(x_1, \dots, x_N) = \sum_{E_1, \dots, E_N} C(E_1, \dots, E_N) \psi_{E_1}(x_1) \dots \psi_{E_N}(x_N)$$

$$= \sum_{n_1, \dots, n_\infty} f(n_1, \dots, n_\infty) \sqrt{\frac{n_1! \dots n_\infty!}{N!}} \sum_{\substack{E_1, \dots, E_N \\ (n_1, \dots, n_\infty)}} \psi_{E_1}(x_1) \dots \psi_{E_N}(x_N)$$

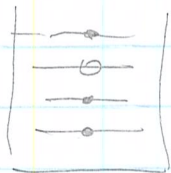
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$$\phi_{n_1, n_2, \dots, n_\infty}(x_1, \dots, x_N) \equiv \sqrt{\frac{n_1! \dots n_\infty!}{N!}} \sum_{\substack{\mathbb{E}_1, \dots, \mathbb{E}_N \\ (n_1, \dots, n_\infty)}} \varphi_{\mathbb{E}_1}(x_1) \dots \varphi_{\mathbb{E}_N}(x_N)$$

$$\mathbb{E}_x: \phi_{2,1,0,\dots,0}(x_1, x_2, x_3) = \frac{1}{\sqrt{3}} \left\{ \varphi_1(x_1) \varphi_1(x_2) \varphi_2(x_3) + \varphi_1(x_1) \varphi_2(x_2) \varphi_1(x_3) + \varphi_2(x_1) \varphi_1(x_2) \varphi_1(x_3) \right\}$$

$\phi_{n_1, n_2, \dots, n_\infty}(x_1, \dots, x_N) \equiv \langle x_1, x_2, \dots, x_N | n_1, n_2, \dots, n_\infty \rangle$
 $|n_1, n_2, \dots, n_\infty\rangle$ base no espaço de Fock.

Fermions:



$$C(\dots \mathbb{E}_i \dots \mathbb{E}_j \dots) = -C(\dots \mathbb{E}_j \dots \mathbb{E}_i \dots)$$

$$\forall \mathbb{E}_i = \mathbb{E}_j \quad C = 0 \Rightarrow n_i = 0, 1.$$

$$\sum_{\mathbb{E}_1, \dots, \mathbb{E}_N} |C(\dots \mathbb{E}_i \dots \mathbb{E}_j \dots)|^2 = 1$$

$$\Rightarrow \bar{C}(n_1, n_2, \dots, n_\infty) \sum_{\substack{\mathbb{E}_1, \mathbb{E}_2, \dots \\ \{n_1, n_2, \dots\}}} 1 = \bar{C}(n_1, \dots, n_\infty) N!$$

$$\bar{C}(n_1, \dots, n_\infty) = C(\mathbb{E}_1, < \mathbb{E}_2, \dots) ; f(n_1, \dots, n_\infty) = \sqrt{N!} \bar{C}(n_1, \dots, n_\infty)$$

$$\phi_{n_1, \dots, n_\infty}(x_1, \dots, x_N) = \frac{1}{\sqrt{N!}} \sum_{\substack{\mathbb{E}_1, \dots, \mathbb{E}_N \\ (n_1, \dots, n_\infty)}} e^{E_1 - E_N} \varphi_{\mathbb{E}_1}(x_1) \dots \varphi_{\mathbb{E}_N}(x_N)$$

$$= \frac{1}{\sqrt{N!}} \begin{vmatrix} \varphi_{\mathbb{E}_1}(x_1) & \dots & \varphi_{\mathbb{E}_1}(x_N) \\ \vdots & & \vdots \\ \varphi_{\mathbb{E}_N}(x_1) & \dots & \varphi_{\mathbb{E}_N}(x_N) \end{vmatrix} \rightarrow \text{determinante de Slater}$$

$|n_1, \dots, n_\infty\rangle \rightarrow$ base no espaço de Fock.

Regras do spin e estatística:

Bosons = spin inteiro = 0, 1, ...

Fermions = spin semi-inteiro = $\frac{1}{2}, \frac{3}{2}, \dots$

$$|1\rangle = \sum_{n_1, \dots, n_\infty} f(n_1, \dots, n_\infty, t) |n_1, \dots, n_\infty\rangle, \rightarrow \text{geral!}$$

Configurações de energia : $r = \{n_1, n_2, \dots, n_\infty\} = \{n_p\}$
e $\{E_p\} = \{E_0, E_1, \dots, E_\infty\}$

$\Rightarrow E_r = E(\{n_p\})$

$T_r = \sum_r = \sum_{\{n_p\}}$

Função de partição : $Z_N = \sum_r e^{-\beta E_r}$ (canônico)

$= \sum_{\substack{\{n_p\} \\ \sum n_p = N}} e^{-\beta E(\{n_p\})} = \sum_{\substack{\{n_p\} \\ \sum n_p = N}} \prod_{l=0}^{\infty} e^{-\beta n_l E_l}$

Gran-canônico:

$Z_G = \sum_{N=0}^{\infty} z^N Z_N(T, V) = \sum_{N=0}^{\infty} \sum_{\substack{\{n_p\} \\ \sum n_p = N}} \prod_{l=0}^{\infty} z^{n_l} e^{-\beta n_l E_l}$; $z = e^{\beta \mu}$

$\sum_{N=0}^{\infty} \sum_{\substack{\{n_p\} \\ \sum n_p = N}} = \sum_N \sum_{\{n_p\}} \delta_{N, \sum n_p} = \sum_{\{n_p\}}$

$\Rightarrow Z_G = \sum_{\{n_p\}} \prod_l (z e^{-\beta E_l})^{n_l} = \prod_l \sum_{n_l} (z e^{-\beta E_l})^{n_l}$

Estatística de Fermi-Dirac : $n_p = 0$ ou 1

$\Rightarrow Z_G = \prod_l (1 + z e^{-\beta E_l})$

Estatística de Bose-Einstein n_p é qualquer.

$\Rightarrow Z_G = \prod_l \frac{1}{1 - z e^{-\beta E_l}}$ se $z e^{-\beta E_l} < 1$

mas $z e^{-\beta E_l} = e^{(\mu - E_l)\beta} < 1 \Rightarrow \mu < E_l \forall l \Rightarrow \mu < E_0$

Assim $Z_G = \prod_l (1 \pm z e^{-\beta E_l})^{\pm 1}$; $\left. \begin{array}{l} + \rightarrow \text{férmions} \\ - \rightarrow \text{bósons.} \end{array} \right\}$

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Max $\Omega = -kT \ln Z_G$, e $\left. \frac{\partial \Omega}{\partial \mu} \right|_{T, V} = -N$

$$\Omega = -kT \ln Z_G = -kT \sum_i \ln (1 \pm z e^{-\beta \epsilon_i})^{\pm 1}$$

$$\Rightarrow N = kT \sum_i \frac{\partial}{\partial \mu} \ln (1 \pm z e^{-\beta \epsilon_i})^{\pm 1}$$

$$= \pm kT \sum_i \frac{\partial z}{\partial \mu} \frac{\partial}{\partial z} \ln (1 \pm z e^{-\beta \epsilon_i})$$

$$\frac{\partial z}{\partial \mu} = \beta e^{\mu/\beta} \Rightarrow N = \pm \sum_i e^{\beta \mu} \frac{(\pm) e^{-\beta \epsilon_i}}{1 \pm z e^{-\beta \epsilon_i}}$$

$$N = \sum_i \frac{e^{\beta(\mu - \epsilon_i)}}{1 \pm e^{\beta(\mu - \epsilon_i)}} = \sum_i \frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1} = \sum_i \langle n_i \rangle$$

\downarrow
n.º de ocupação
média

- (+) e $\forall \mu$ p/ férmions
- (-) e $\mu \leq \epsilon_0$ p/ bósons.

Limite clássico: Maxwell-Boltzmann.

Neste caso $\langle n_i \rangle_{FD} = \langle n_i \rangle_{BE} \Rightarrow z \ll 1$

$$e \langle n_i \rangle_{MB} = e^{-\beta(\epsilon_i - \mu)}$$

$$Z_G = e^{\sum_i \ln (1 \pm z e^{-\beta \epsilon_i})^{\pm 1}} = e^{\pm \sum_i \ln (1 \pm z e^{-\beta \epsilon_i})}$$

$$= e^{\pm \sum_i z e^{-\beta \epsilon_i}} = e^{\pm z \sum_i e^{-\beta \epsilon_i}}$$

$$\Rightarrow N = - \frac{\partial \Omega}{\partial \mu} = kT \cdot \frac{\partial}{\partial \mu} \ln \Omega$$

$$= kT \frac{\partial z}{\partial \mu} \frac{\partial}{\partial z} \left(z \sum_i e^{-\beta \epsilon_i} \right) = \sum_i e^{-\beta(\epsilon_i - \mu)}$$

$$\Rightarrow \langle n_e \rangle_{MB} = e^{-\beta(\epsilon_e - \mu)}$$

Voltamos à Z_G : $Z_G = e^{z \sum e^{-\beta \epsilon_i}}$

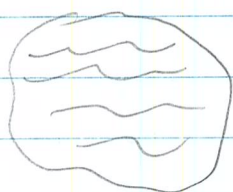
$$= \sum_{N=0}^{\infty} \frac{z^N}{N!} \left(\sum_i e^{-\beta \epsilon_i} \right)^N \quad N = \sum_i n_i$$

$$\Rightarrow Z_N = \frac{\left(\sum_i e^{-\beta \epsilon_i} \right)^N}{N!} = \sum_{\{n_i\}} \frac{1}{n_0! n_1! \dots n_\infty!} e^{-\beta \epsilon_0 n_0 - \beta \epsilon_1 n_1 - \dots}$$

\Rightarrow justificativa do $N!$ presente no caso clássico.

Potencial químico e relatividade

Nº de fótons é variável porque equilíbrio do gás a temperatura T depende de emissão e absorção.



$F(T, V, N)$ é mínimo p/ $T = \text{const}$
 $V = \text{const}$, mas N não é constante

$$\Rightarrow dF = -S dT - p dV + \mu dN = 0$$

$$\Rightarrow \mu = \left. \frac{\partial F}{\partial N} \right|_{T, V} = 0!$$

Se há lei de conservação: $e^{(+)} + e^{(-)} \rightarrow 2\gamma$

$$N^{(-)} - N^{(+)} = N_0 \rightarrow \text{fixo}$$

$$F(T, V, N_0; N^{(-)}) = F(T, V; N^{(-)}, N^{(+)} = -N_0 + N^{(-)})$$

$$\Rightarrow \left. \frac{\partial F}{\partial N^{(-)}} \right|_{N_0} = \left. \frac{\partial F}{\partial N^{(-)}} \right|_{N^{(+)}} + \left. \frac{\partial F}{\partial N^{(+)}} \right|_{N^{(-)}} \frac{\partial N^{(+)}}{\partial N^{(-)}} = \mu^{(-)} + \mu^{(+)} = 0$$

$$\Rightarrow \mu^{(+)} = -\mu^{(-)}$$

$$\text{ou } e^{(+)} + e^{(-)} \rightarrow 2\gamma$$

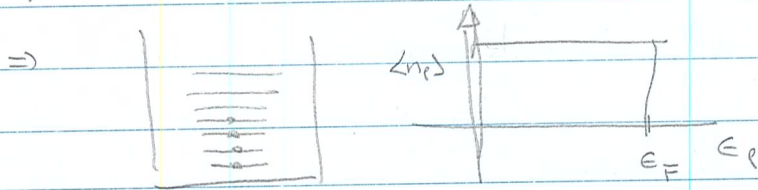
$$\sum \nu_i \mu_i = 0 \Rightarrow \nu^{(+)} \mu^{(+)} + \nu^{(-)} \mu^{(-)} + \nu^{(\gamma)} \mu^{(\gamma)} = 0$$

Energia relativística: $E_p = \sqrt{p^2 c^2 + m^2 c^4}$

$$\langle n_p \rangle = \frac{1}{e^{\beta(E_p - \mu)} + 1} \quad e \quad \mu < E_0 = mc^2$$

Gás ideal de Fermions:

A $T=0$ E é mínima.



$$n_p = \begin{cases} 1 & \text{se } E_p < E_F \\ 0 & \text{se } E_p > E_F \end{cases} \Rightarrow n_p = \theta(E_F - E_p)$$

De fato, lim $\beta \rightarrow \infty \frac{1}{e^{\beta(E_p - \mu)} + 1} = \theta(\mu - E_p)$

$$\Rightarrow \mu(T=0) = E_F$$

Mas $\mu(T=0) = \frac{\partial E}{\partial N} = E(N) - E(N-1) = E_F$

Se o gás tem massa m e spin s (semi-inteiro) $\{\vec{p}, s_z\}$ $g = (2s+1)$

$$\sum_1 = \sum_{\vec{p}, s_z} = g \frac{V}{h^3} \int d^3 p \quad \text{pass } \sum \rightarrow \frac{V}{(2\pi)^3} \int d^3 k$$

$$N = g \frac{V}{h^3} \int_{p \leq p_F} d^3 p = g \frac{V}{h^3} \int_0^{p_F} 4\pi p^2 dp = g \frac{V}{h^3} \frac{4\pi}{3} p_F^3$$